Physics 222: Modern Physics for Engineers Spring 2021

Instructor: Hans Schuessler Website: <u>http://sibor.physics.tamu.edu</u> Office: 442 MPHY

Email: schuessler@physics.tamu.edu

Course objectives: To learn the physics of the 20th century Course outcomes: Know the basic laws of relativity, quantum and atomic physics, nuclear physics ad solid state physics.

Text: Modern Physics for Scientists and Engineers by S. Thornton and A. Rex, 5th Edition, ISBN: 9781337919456

CHAPTER 1

"The eternal mystery of the world is its comprehensibility." "Anyone who has never made a mistake has never tried anything new."

A. Einstein

- 1.1 Classical Physics of the 1890s
- 1.2 The Kinetic Theory of Gases
- 1.3 Waves and Particles
- 1.4 Conservation Laws and Fundamental Forces
- 1.5 The Atomic Theory of Matter
- 1.6 Unresolved Questions of 1895 and New Horizons

1.1: Classical Physics of the 1890s

- Mechanics
- Electromagnetism
- Thermodynamics



Triumph of Classical Physics: The Conservation Laws

- Conservation of energy: The total sum of energy (in all its forms) is conserved in all interactions.
- Conservation of linear momentum: In the absence of external forces, linear momentum is conserved in all interactions.
- Conservation of angular momentum: In the absence of external torque, angular momentum is conserved in all interactions.
- Conservation of charge: Electric charge is conserved in all interactions.

Mechanics



- Galileo Galilei (1564-1642)
- Great experimentalist
- Principle of inertia
- Established experimental foundations

Kinematics equations for constant acceleration

General formulas:

$$v = dx / dt$$

$$x = x_0 + v_0 t + \frac{1}{2} a t^2$$

$$a = dv / dt = d^2 x / dt^2$$

$$v^2 = v_0^2 + 2a(x - x_0)$$

$$x = x_0 + (\frac{v_0 + v}{2})t$$

$$\frac{\text{hotion with const acceleration}}{a = \frac{\upsilon_2 - \upsilon_1}{t_2 - t_1} = \frac{\upsilon - \upsilon_0}{t_2 - 0} \qquad (\underline{\upsilon = \upsilon_0 + at}) \\ = \frac{\upsilon_1 - \upsilon_1}{t_2 - t_1} = \frac{\upsilon_1 - \upsilon_0}{t_2 - 0} \qquad (\underline{\upsilon = \upsilon_0 + at}) \\ = \frac{\upsilon_0 + \upsilon_1}{t_2 - t_1} = \frac{\upsilon_1 - \upsilon_0}{t_2 - 0} \qquad (\underline{\upsilon = \upsilon_0 + \frac{1}{2}at}) \\ = \frac{\varepsilon_1 - \varepsilon_1}{t_2 - t_1} = \frac{\varepsilon_1 - \varepsilon_0}{t_2 - 0} \qquad (\underline{\varepsilon_1 - \varepsilon_0})^2 \qquad (\underline{\varepsilon_1 - \varepsilon_0})^2 \\ = \frac{\varepsilon_1 - \varepsilon_1}{t_2 - t_1} = \frac{\varepsilon_1}{t_2 - t_1}$$

Testing Kinetics for a=9.80m/s²



All objects fall with the same constant acceleration!!



Isaac Newton (1642-1727)

Three laws describing the relationship between mass and acceleration.



- Newton's first law (law of inertia): An object in motion with a constant velocity will continue in motion unless acted upon by some net external force.
- Newton's second law: Introduces force (F) as responsible for the change in linear momentum (p):

$$\vec{F} = m\vec{a}$$
 or $\vec{F} = \frac{d\vec{p}}{dt}$

Newton's third law (law of action and reaction): The force exerted by body 1 on body 2 is equal in magnitude and opposite in direction to the force that body 2 exerts on body 1.

$$\vec{F}_{21} = -\vec{F}_{12}$$

Inertial Frames K and K'



- K is at rest and K' is moving with constant velocity
- Axes are parallel
- K and K' are said to be INERTIAL COORDINATE SYSTEMS

The Galilean Transformation

For a point P

- In system K: P = (x, y, z, t)
- In system K': P = (x', y', z', t')



pulling a sled, Michelangelo's assistant



For forward motion: $F_{AG} > F_{AS}$ $F_{SA} > F_{SG}$



Newton's Law of Gravitation

$$F_g = \frac{Gm_1m_2}{r^2}$$



G=gravitational constant = $6.673(10) \times 10^{-11} Nm^2 / kg^2$

Note: The weight ω of a body of mass m on the earth's surface with radius R_E is $\omega = mg = \frac{Gm_E \cdot m}{R_E^2}$ or $g = \frac{Gm_E}{R_E^2}$

Cavendish balance



Henry Cavendish(1798) announced that he has weighted the earth.

$$a_{\text{porth}} = 9.81 \text{ M}_{\text{S}^2}$$

$$a_{\text{moory}} = 1.67 \text{ M}_{\text{S}^2}$$

$$a_{\text{sum}} = 274 \text{ M}_{\text{S}^2}$$

 $a_{gr} = g = G \frac{M}{R^2}$

The Foucault pendulum in Aggieland

What does it show?? Seeing is believing T=10 sec how long is it??



Electromagnetism: 18th-19th centuries

Contributions made by:

- Coulomb (1736-1806)
- Oersted (1777-1851)
- Young (1773-1829)
- Ampère (1775-1836)
- Faraday (1791-1867)
- Henry (1797-1878)
- Maxwell (1831-1879)
- Hertz (1857-1894)















Culminates in Maxwell's Equations

Gauss' s law (Φ_E):
 (electric field)

(magnetic field)

Gauss' s law (Φ_B) :

 $\oint \vec{E} \cdot d\vec{A} = \frac{q}{\varepsilon_0}$

$$\oint \vec{B} \cdot d\vec{A} = 0$$

Gauss (1777 –1855)

- Faraday' s law:
- Ampère' s law: (Generalized)
- Lorentz law: (force)

$$\oint \vec{E} \cdot d\vec{s} = -\frac{d\Phi_B}{dt}$$

$$\oint \vec{B} \cdot d\vec{s} = \mu_0 \varepsilon_0 \frac{d\Phi_E}{dt} + \mu_0 I$$

 $\overline{F} = q\overline{E} + q\nabla \times \overline{B}$

Thermodynamics

Contributions made by:

- Benjamin Thompson (1753-1814) (frictional heat) (Count Rumford)
- Sadi Carnot (1796-1832) (heat engine and Carnot cycle)
- James Joule (1818-1889) (mechanical equivalent of heat))
- Rudolf Clausius (1822-1888) (heat can never pass from a colder to a warmer body without some other change)
- William Thompson (1824-1907) (Lord Kelvin) (proposed absolute temperature scale)

Primary Results

- Deals with temperature, heat, work, and the internal energy of systems
- Introduces thermal equilibrium
- The first law establishes heat as energy and expresses conservation of energy
- Introduces the concept of internal energy and considers temperature as a measure of internal energy
- Introduces limitations of the energy processes that can or cannot take place

The Laws of Thermodynamics

First law: The change in the internal energy ΔU of a system is equal to the heat Q added to a system plus the work W done on the system

 $\Delta U = Q + W$

- Second law: It is not possible to convert heat completely into work without some other change taking place.
- **The "zeroth" law**: Two systems in thermal equilibrium with a third system are in thermal equilibrium with each other.
- Third law: It is not possible to achieve an absolute zero temperature.

1.2: The Kinetic Theory of Gases

Contributions made by:

- Robert Boyle (1627-1691)
- Jacques Charles (1746-1823)
- Joseph Louis Gay-Lussac (1778-1823)
- Culminates in the ideal gas equation for n moles of a "simple" gas:

$$pV = nRT$$

(where R is the ideal gas constant, 8.31 J/(mol · K)

Additional Contributions

- Amedeo Avogadro (1776-1856) (number of molecules in a mole and their weights)
- John Dalton (1766-1844) (atomic theory of elements)
- Daniel Bernoulli (1700-1782) (kinetic theory of gases)
- Ludwig Boltzmann (1844-1906) (kinetic theory of gases, entropy as log() of probability)
- James Clerk Maxwell (1831-1879) (velocity distribution)
- J. Willard Gibbs (1939-1903) (thermodynamic and chemical potentials)

Primary Results

- Average molecular kinetic energy directly related to absolute temperature
- Internal energy U directly related to the average molecular kinetic energy
- Internal energy equally distributed among the number of degrees of freedom (f) of the system containing n moles of substance

$$U = nN_A \langle K \rangle = \frac{f}{2} nRT$$

$$(N_A = 6.022 \times 10^{23} \text{ mol}^{-1})$$

Avogadro's number: number of molecules in 1 mole)

Other Primary Results

2. Maxwell derives a relation for the molecular speed distribution f(v):

$$f(v) = 4\pi N \left(\frac{m}{2\pi kT}\right)^{3/2} v^2 e^{-mv^2/2kT}$$

3. Boltzmann contributes to determine the *root-mean-square* of the molecular speed

$$v_{rms} = \sqrt{\langle v^2 \rangle} = \sqrt{\frac{3kT}{m}}$$

Thus relating energy to the temperature for an ideal gas

Molecular speeds in an ideal gas

$$K_{av}(molecule) = \frac{1}{2}mv_{ev}^{2} = \frac{3}{2}kT \left[v_{rms} = \sqrt{v_{av}^{2}} = \sqrt{\frac{3\kappa\Gamma}{m}} = \sqrt{\frac{3\kappa\Gamma}{m}} \right]$$

$$(onsider 0_{2} as an ideal goas at 27°C. Find K_{av} and K_{4r} (27°C = 300 \text{ K})$$

$$k_{av} = \frac{1}{2}mv_{av}^{2} = \frac{3}{2}kT = \frac{3}{2}(1.39 \times 10^{-23})(300) = 6.21\times10^{-21}$$

$$k_{av} = \frac{3}{2}\kappaT = \frac{3}{2}(1.39\times10^{-23})(300) = 6.21\times10^{-21}$$

$$k_{tr} = \frac{3}{2} RT = \frac{3}{2} (imal) \frac{8.3}{5} \cdot \frac{300k}{5} = \frac{3740}{5}$$

Number of molecules



1.3: Waves and Particles

Two ways in which energy is transported:

- 1) Point mass interaction: transfers of momentum and kinetic energy: *particles*
- Extended regions wherein energy transfers
 by way of vibrations and rotations are
 observed: waves

Particles vs. Waves

- Two distinct phenomena describing physical interactions
 - Particles in the form of point masses and waves in the form of perturbation in a mass distribution, i.e., a material medium
 - The distinctions are observationally quite clear; however, not so for the case of visible light
 - Thus by the 17th century begins the major disagreement concerning the nature of light

The Nature of Light

Contributions made by:

- Isaac Newton (1642-1742)
- Christian Huygens (1629 1695)
- Thomas Young (1773 -1829)
- Augustin Fresnel (1788 1829)

The Nature of Light

- Newton suggested the corpuscular (particle) theory
 - Particles of light travel in straight lines or rays
 - Explained sharp shadows
 - Explained reflection and refraction

The Nature of Light

Christian Huygens promoted the wave theory

- Light propagates as a wave of concentric circles from the point of origin
- Explained reflection and refraction
- Did not explain sharp shadows

The Wave Theory Advances...

- Contributions by Huygens, Young, Fresnel and Maxwell
- Double-slit interference patterns
- Refraction of light from a vacuum to a medium
- Light is an electromagnetic phenomenon
- Establishes that light propagates as a wave

The Electromagnetic Spectrum

- Visible light covers only a small range of the total electromagnetic spectrum
- All electromagnetic waves travel in a vacuum with a speed c given by:

$$c = \frac{1}{\sqrt{\mu_0 \varepsilon_0}} = \lambda f$$

(where μ_0 and ε_0 are the respective permeability and permittivity of "free" space)

1.4: Conservation Laws and Fundamental Forces

- Recall the fundamental conservation laws:
 - Conservation of energy
 - Conservation of linear momentum
 - Conservation of angular momentum
 - Conservation of electric charge
- Later we will establish the conservation of mass as part of the conservation of energy

Modern Results

- In addition to the classical conservation laws, two modern results will include:
 - The conservation of baryons and leptons
 - The fundamental invariance principles for time reversal, distance, and parity

Also in the Modern Context...

The three fundamental forces are introduced

Gravitational:
$$\vec{F}_g = -G \frac{m_1 m_2}{r^2} \hat{r}$$

- Electroweak
 - **Weak:** Responsible for nuclear beta decay and effective only
 - over distances of ~10⁻¹⁵ m **Electromagnetic**: $\vec{F}_C = \frac{1}{4\pi\varepsilon_0} \frac{q_1q_2}{r^2} \hat{r}$ (Coulomb force)
- **Strong**: Responsible for "holding" the nucleus together and effective less than $\sim 10^{-15}$ m

Unification

- Neutrons and protons are composed of quarks, which have the color force acting between them
- Grand Unified Theory (GUT) attempts to unify electroweak and strong forces
 - String theory is one of these
 - They have yet to be verified experimentally

Unification of Forces

- Maxwell unified the electric and magnetic forces as fundamentally the same force; now referred to as the electromagnetic force
- In the 1970's Glashow, Weinberg, and Salam proposed the equivalence of the electromagnetic and the weak forces (at high energy); now referred to as the electroweak interaction

Goal: Unification of All Forces into a Single Force



1.5: The Atomic Theory of Matter

- Initiated by Democritus and Leucippus (~450 B.C.) (first to us the Greek *atomos*, meaning "indivisible")
- In addition to fundamental contributions by Boyle, Charles, and Gay-Lussac, Proust (1754 – 1826) proposes the law of definite proportions
- Dalton advances the atomic theory of matter to explain the law of definite proportions
- Avogadro proposes that all gases at the same temperature, pressure, and volume contain the same number of molecules (atoms); viz. 6.02 × 10²³ atoms per mole
- Cannizzaro (1826 1910) makes the distinction between atoms and molecules advancing the ideas of Avogadro.

Further Advances in Atomic Theory

- Maxwell derives the speed distribution of atoms in a gas
- Robert Brown (1753 1858) observes microscopic "random" motion of suspended grains of pollen in water
- Einstein in the 20th century explains this random motion using atomic theory

Opposition to the Theory

- Ernst Mach (1838 1916) opposes the theory on the basis of logical positivism, i.e., atoms being *"unseen" place into question their reality*
- Wilhelm Ostwald (1853 1932) supports this premise but on experimental results of radioactivity, discrete spectral lines, and the formation of molecular structures

Overwhelming Evidence for Existence of Atoms

- Max Planck (1858 1947) advances the concept to explain blackbody radiation by use of submicroscopic "quanta"
- Boltzmann requires existence of atoms for his advances in statistical mechanics
- Albert Einstein (1879 1955) uses molecules to explain Brownian motion and determines the approximate value of their size and mass
- Jean Perrin (1870 1942) experimentally verifies Einstein's predictions

1.6: Unresolved Questions of 1895 and New Horizons

- The atomic theory controversy raises fundamental questions
 - It was not universally accepted
 - The constitutes (if any) of atoms became a significant question
 - The structure of matter remained unknown with certainty

Further Complications

Three fundamental problems:

- The question of the existence of an electromagnetic medium
- The problem of observed differences in the electric and magnetic field between stationary and moving reference systems
- The failure of classical physics to explain blackbody radiation

Additional Discoveries Contribute to the Complications

- Discovery of x-rays
- Discovery of radioactivity
- Discovery of the electron
- Discovery of the Zeeman effect

The Beginnings of Modern Physics

- These new discoveries and the many resulting complications required a revision of the fundamental physical assumptions that culminated in the huge successes of physics
- To this end, the introduction of the modern theory of relativity and quantum mechanics becomes the starting point of this most fascinating revision.

Four forces video